DISCOVERIES OF ROTATING RADIO TRANSIENTS

IN THE 350 MHZ GREEN BANK TELESCOPE DRIFT-SCAN SURVEY

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DRIFT-SCAN COLLABORATION

- Anne Archibald (McGill)
- Aaron Berndsen (UBC)
- Jason Boyles (WVU)
- Fernando Cardoso (WVU)
- Angus Cherry (UBC)
- Jason Hessels (ASTRON)
- Vicky Kaspi (McGill)
- Vlad Kondratiev (ASTRON)
- Joeri van Leeuwen (ASTRON)

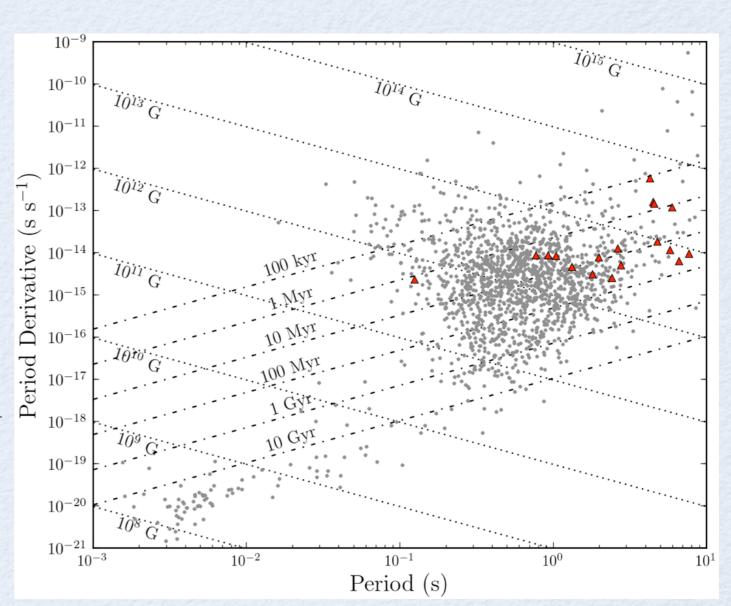
- Duncan Lorimer (WVU)
- Ryan Lynch (McGill)
- Maura McLaughlin (WVU)
- Christie McPhee (UBC)
- Tim Pennucci (UVA)
- Scott Ransom (NRAO)
- Mallory Roberts (Eureka)
- Ingrid Stairs (UBC)

WHAT'S A RRAT?

- Definitions vary...
 - A pulsar only detectable by its single pulses
 - Pulses not detectable every period; up to ~1000 periods of separation between pulses
 - Won't be seen in periodicity searches

MOITAVITON

- Why search for RRATs?
- Known RRAT population very small
- Don't yet know where fit in with pulsar population
 - Evolutionary link?
- RRATs may outnumber regular pulsars!



RRAT parameters taken from The RRATalog (McLaughlin).

GBT DRIFT-SCAN SURVEY

- How do we search for RRATs?
- GBT 350 MHz Drift-Scan Survey
 - 30,000 140-s pointings, 10347
 sq deg
- Produce diagnostic single pulse plots as part of survey processing (PRESTO)

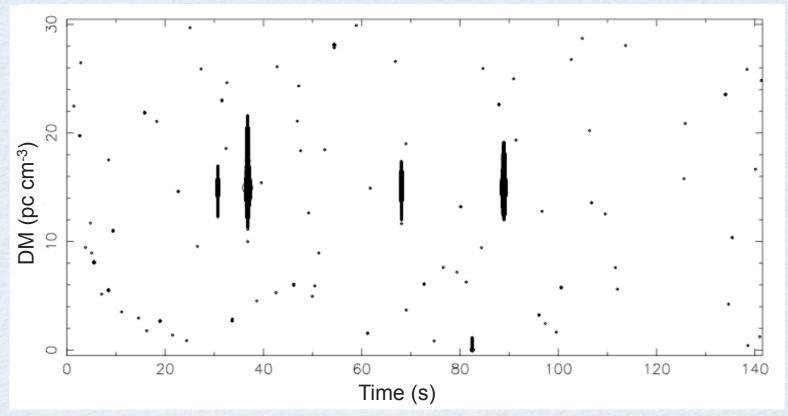


Image credit: NRAO

Manually examine plots

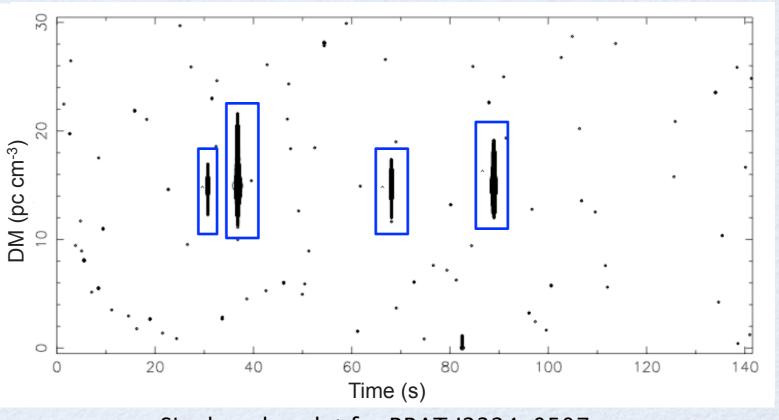
- Why automate the search for RRATs?
- 30,000+ pointings in GBT drift-scan survey = 30,000+ single pulse plots to manually examine = *tedious*!
- Developed a post-processing identification algorithm for single pulse candidates: distinguish astrophysical pulses from RFI
- Universal: can be applied to future surveys

- Initially: single pulse candidate events for every beam
 - dot size ~ S/N
- Idea: group together events which may belong to the same pulse, based on proximity in t, DM



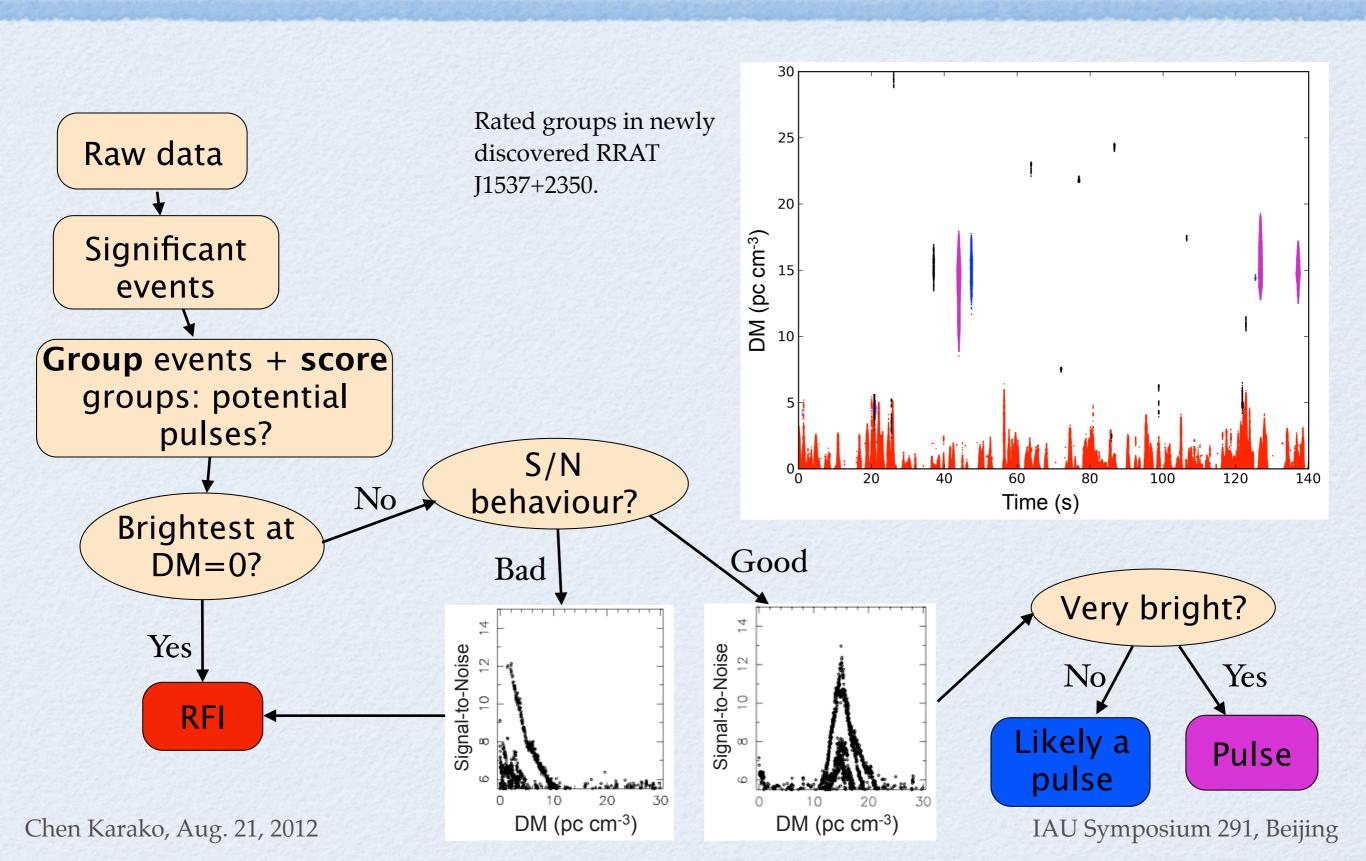
Single pulse plot for RRAT J2324-0507.

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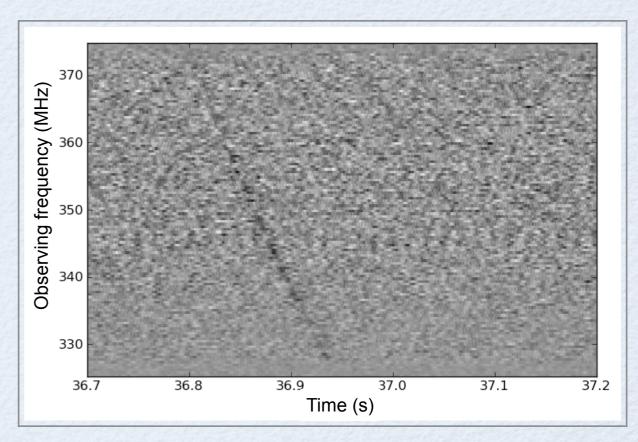


Single pulse plot for RRAT J2324-0507.

• then: **score** groups to distinguish astrophysical pulses from RFI



- Finally, examine flagged beams
- Go back to raw data to check for dispersion relation ("waterfall plots")



RRAT J2324-0507

RESULTS

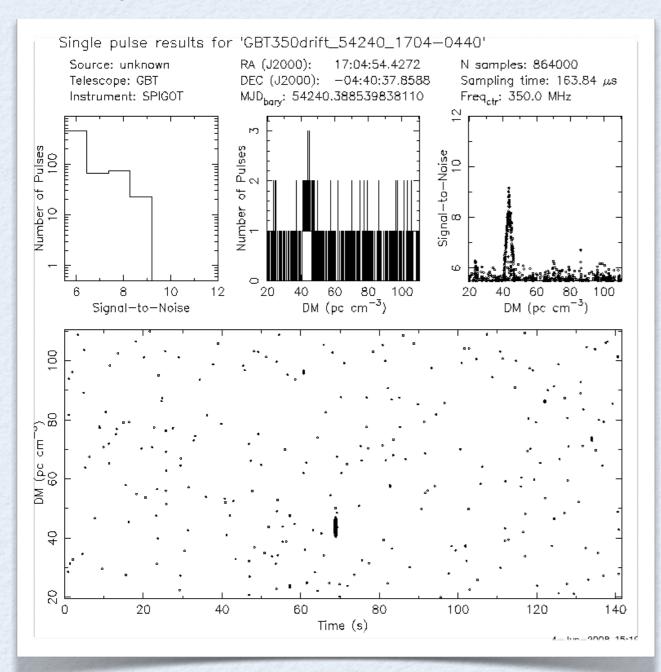
- 34 RRAT candidates found overall in drift-scan survey
 - vs. 31 regular pulsars (!)
- Of the 8 followed up, 7 have been confirmed
- Hope to begin timing these sources soon
- Remaining candidates await confirmation observations

Table 1. Properties of New Drift-Scan RRATs									
	•	y Position							
Name	R.A. (J2000)	Dec. (J2000)	$_{\mathrm{DM}}$	# Pulses detected	Burst rate*	P^*			
	(h m s)	(° ′ ″)	$(pc cm^{-3})$	(discovery)	(/hr)	(s)			
RRAT 1704-0440	17:04:54.43	-04:40:37.9	43 ± 1	1		0.238			
RRAT 1537+2350	15:37:47.36	+23:50:51.1	15 ± 1	5	77 ± 14	3.45			
RRAT 1944-1017	19:44:08.58	-10:17:07.2	31 ± 1	9	77 ± 14	0.409			
RRAT 2324-0507	23:24:22.17	-05:07:36.0	15 ± 1	4	46 ± 9	0.869			
RRAT 2006+2021	20:06:54.79	+20:21:05.8	66 ± 1	3		4.63			
RRAT 0206-0443	02:06:11.00	-04:43:33.4	15 ± 1	1					
RRAT 0447-0442	04:47:03.79	-04:42:15.2	30 ± 1	4		4.38			
RRAT 0513-0417	05:13:20.72	-04:17:53.4	19 ± 1	1					
RRAT 0526-1908	05:26:04.86	-19:08:48.8	26 ± 2	1					
RRAT 0544-0309	05:44:46.42	-03:09:44.9	68 ± 2	3					
RRAT 0606-0129	06:06:37.50	-01:29:23.9	17 ± 3	1					
RRAT 0614-0328	06:14:37.82	-03:28:44.9	18±1	5		0.136			
RRAT 0645-0128	06:45:39.55	-01:28:58.1	15 ± 2	2					
RRAT 0706-1058	07:06:43.16	-10:58:20.3	19 ± 1	1					
RRAT 0719-1907	07:19:38.34	-19:07:30.2	27 ± 1	3					
RRAT 0807-1057	08:07:02.66	-10:57:41.4	18±1	2					
RRAT 0808-0746	08:08:43.30	-07:46:58.9	36 ± 1	1					
RRAT 0858-0746	08:58:23.86	-07:46:31.4	40 ± 3	4		0.884			
RRAT 0957-0617	09:57:09.05	-06:17:20.5	27 ± 1	7		1.72			
RRAT 1059-0102	10:59:31.93	-01:02:01.6	18±1	1					
RRAT 1132+2455	11:32:09.18	+24:55:58.7	24 ± 1	3		1.00			
RRAT 1222-1806	12:22:25.02	-18:06:27.5	17 ± 1	1					
RRAT 1236-1806	12:36:36.40	-18:06:28.1	18 ± 1	1					
RRAT 1253-0101	12:53:04.55	-01:01:56.4	170 ± 5	1					
RRAT 1332-0326	13:32:14.78	-03:26:26.3	27 ± 1	2					
RRAT 1336-2034	13:36:28.21	-20:34:21.8	19 ± 1	3		0.184			
RRAT 1610-0128	16:10:58.44	-01:28:04.0	27 ± 1	1					
RRAT 1914-1129	19:14:46.31	-11:29:33.5	90±1	1					
RRAT 2335-0443	23:35:51.41	-04:43:50.3	15 ± 1	1					
*Note the period may be an integer fraction of the tabulated value.									

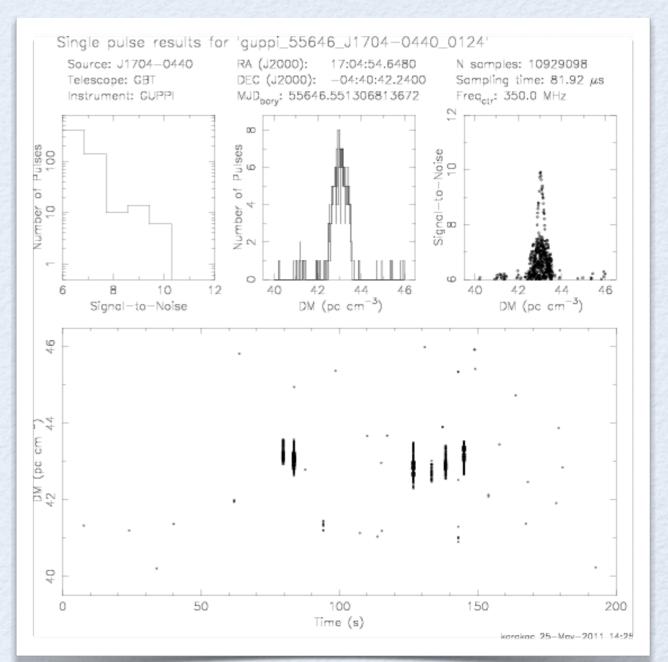
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DISCOVERY & CONFIRMATION

• J1704-0440, P=0.238 s

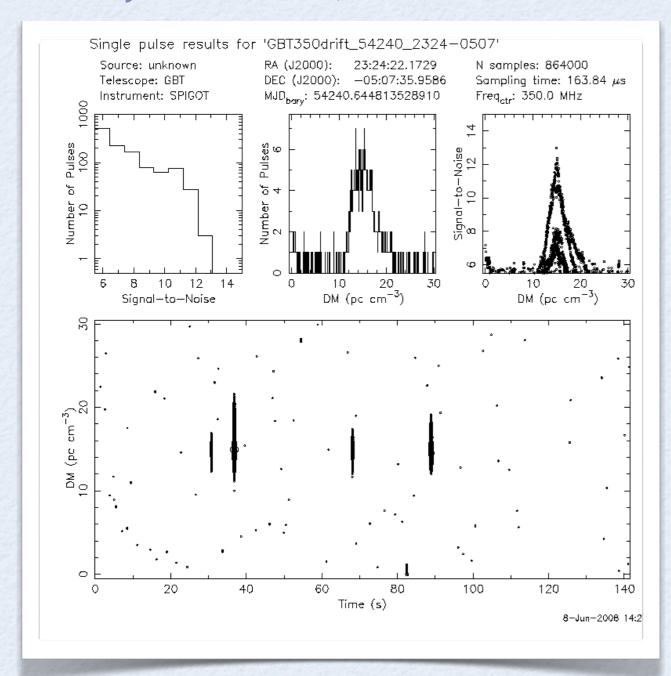


✓ CONFIRMED

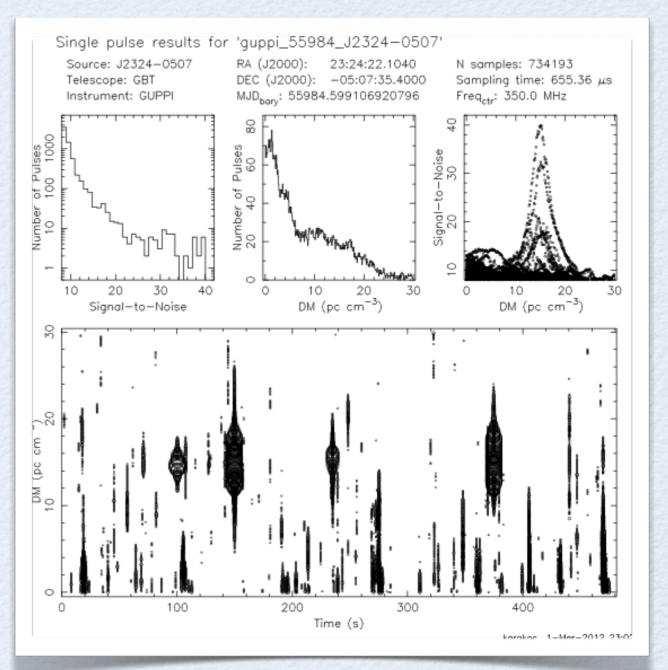


DISCOVERY & CONFIRMATION

• J2324-0507, P=0.869 s

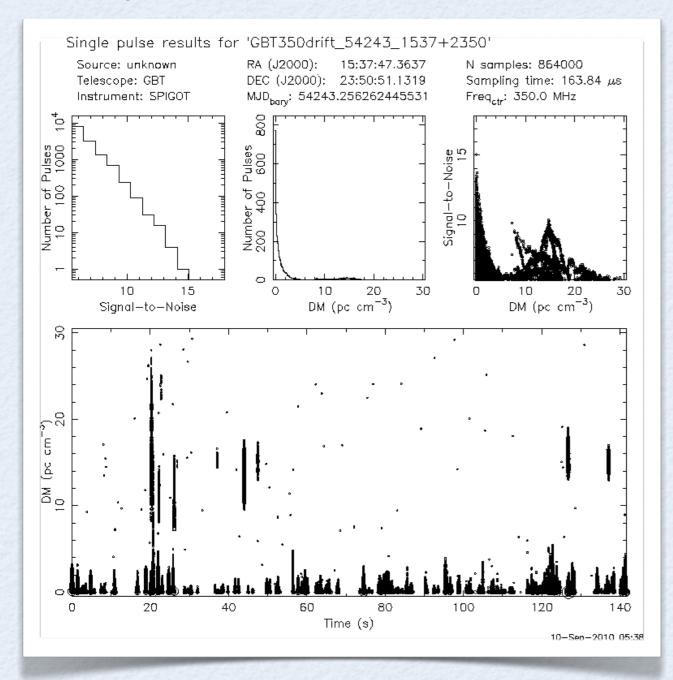


√ CONFIRMED

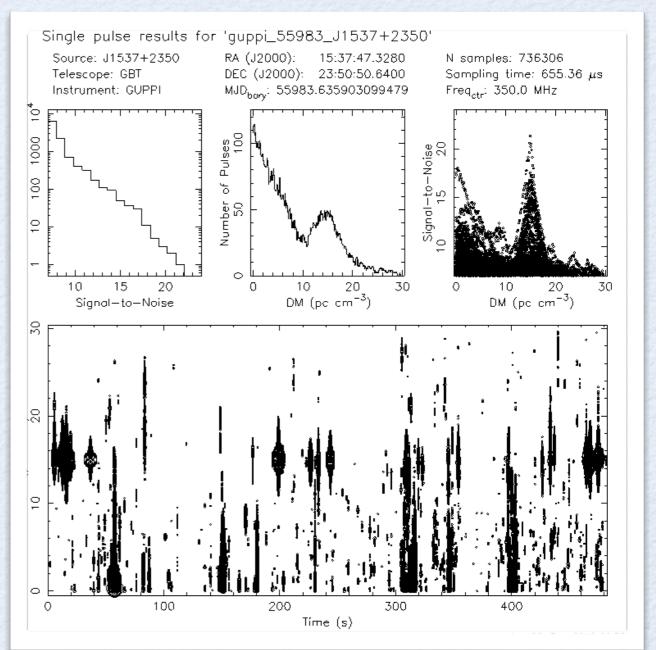


DISCOVERY & CONFIRMATION

• J1537+2350, P=3.45 s



√ CONFIRMED



LOOKING AHEAD...

- Timing of confirmed sources
- Follow up of remaining 26 candidates
- Improvements to algorithm will allow to identify weaker pulses (there may still be more RRATs in the data!)
- Making code more robust to RFI will further reduce time spent manually examining flagged plots
- Implement code on other surveys, eg. GBNCC